

**BSc/MSci EXAMINATION**

PHY-302 Nuclear Physics And Astrophysics

Time Allowed: 2 hours 15 minutes

Date: 16<sup>th</sup> May 2007

Time: 10:00

Answer ALL questions in section A. Answer ONLY TWO questions from section B. Section A carries 40 marks, each question in section B carries 30 marks. An indicative marking-scheme is shown in square brackets [ ] after each part of a question.

COMPLETE ALL ROUGH WORKINGS IN THE ANSWER BOOK AND CROSS THROUGH ANY WORK WHICH IS NOT TO BE ASSESSED.

NUMERIC CALCULATORS ARE PERMITTED IN THIS EXAMINATION.

**Data**

Electronic charge	e	$1.602 \times 10^{-19} \text{ C}$	
Electron mass	$m_e$	$0.511003 \text{ MeV}/c^2$	$5.485803 \times 10^{-4} \text{ u}$
Proton mass	$m_p$	$938.280 \text{ MeV}/c^2$	$1.00727647 \text{ u}$
Neutron mass	$m_n$	$939.573 \text{ MeV}/c^2$	$1.00866501 \text{ u}$
Atomic mass unit	u	$931.502 \text{ MeV}/c^2$	
Avogadro constant	$N_a$	$6.022045 \times 10^{23} \text{ g mol}^{-1}$	
electron volt	eV	$1.602189 \times 10^{-19} \text{ J}$	
Boltzmann constant	k	$8.6174 \times 10^{-11} \text{ MeV/K}$	

Mass	$^{236}_{92}\text{U}$	236.045568 u
Mass	$^{235}_{92}\text{U}$	235.043924 u

$$e^2/(4\pi\epsilon_0) = 1.44 \text{ MeV fm}$$

If needed assume the ordering of nuclear shells to be

$1s_{1/2}$ ;  $1p_{3/2}$ ;  $1p_{1/2}$ ;  $1d_{5/2}$ ;  $1d_{3/2}$ ;  $2s_{1/2}$ ;  $1f_{7/2}$ ;  $1f_{5/2}$ ;  $2p_{3/2}$ ;  $2p_{1/2}$ ;  $1g_{9/2}$ ;  $1g_{7/2}$ ;  $2d_{5/2}$ ;  $2d_{3/2}$ ;  $1h_{11/2}$ ;  $3s_{1/2}$ ;  $1h_{9/2}$ ;

**YOU ARE NOT PERMITTED TO START READING THIS QUESTION PAPER UNTIL INSTRUCTED TO DO SO BY AN INVIGILATOR.**

**SECTION A: Answer all questions in this section**

- A1. Describe in words what is meant by the thermal reproduction factor  $k_{\infty}$ . Give the equation defining  $k_{\infty}$  and describe each of the four terms. [5]
- A2. Give an approximate value in metres for the nuclear radius of  $^{36}\text{Cl}$  showing how you arrived at your answer. Estimate the atomic radius of  $^{36}\text{Cl}$ ? [4]
- A3. Describe in words what is meant by the cross section of a nuclear process. [2]
- A4. Write down the three reaction equations for  $\beta^+$ ,  $\beta^-$ , and  $\epsilon$  (electron capture) processes. [3]
- A5. Describe briefly the difference between the pp and CNO cycles of hydrogen burning in stars. Reaction equations and Q values are not required. Does the CNO cycle dominate in hotter or colder stars? Explain why. [4]
- A6. Give the formula used to calculate the mass of a nucleus  $M(Z,A)$  given the binding energy B. Calculate the mass of Mg which has  $Z=12$  and  $A=26$  if the binding energy is  $218 \text{ MeV}/c^2$ . [4]
- A7. Explain why alpha particles have a much shorter range in air than beta particles of similar energy. [2]
- A8. Define the Q of a nuclear reaction in terms of nuclear masses. By considering the decay reaction  $A \rightarrow B + C$  derive the Q value in terms of initial and final kinetic energies. What interpretation should be given to a reaction which has a negative Q value? [4]
- A9. Describe briefly what is meant by the r and s processes of nucleosynthesis. What are their main distinguishing features? What conditions are required for these two processes? [4]
- A10. Considering a hot gas of  $^{10}_5\text{B}$  nuclei, calculate classically the temperature required to overcome the Coulomb barrier and the energy release in the fusion reaction. Why are stellar fusion reactions observed to occur at much lower temperatures? [4]
- A11. Deduce from the shell model the spin and parity assignments for the ground states of the following nuclei:
- a)  $^{11}_5\text{B}$
  - b)  $^{17}_9\text{F}$  [4]

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**SECTION B: Answer two questions from this section****B1)**

- a) For a sample of  $N$  nuclei present at time  $t$ , the number of nuclei  $dN$  decaying in a time  $dt$  is given by a differential equation characterising the decay constant  $\lambda$ . Write down this equation and its general solution for  $N(t)$ . [3]
- b) Give equations for the mean lifetime  $\tau$ , and the half life  $t_{1/2}$  in terms of the decay constant  $\lambda$ . Describe in words the meaning of half life. Sketch a graph of  $N$  vs  $t$  for a sample of  $10^{24}$  nuclei which undergo radioactive decay with a decay constant of  $0.005 \text{ days}^{-1}$ . Show on the graph the number of nuclei left after a time  $t = \tau$  and  $t = t_{1/2}$ . [5]
- c)  $^{212}_{83}\text{Bi}$  undergoes  $\alpha$  decay to Titanium (Ti) and  $\beta^-$  decay to Polonium (Po). Write down the equations for both of these reactions giving the  $A$  and  $Z$  values. If the half life of  $\alpha$  decay is  $14.71 \text{ s}$  and the total decay rate is  $0.536 \text{ s}^{-1}$  what is the half life of the  $\beta^-$  decay mode? [5]
- d)  $^{210}_{83}\text{Bi}$  has a mean lifetime of  $7.2 \text{ days}$  and decays via  $\beta^-$ -particle emission to  $^{210}\text{Po}$  which in turn decays via  $\alpha$ -particle emission to  $^{206}\text{Pb}$  with a mean lifetime  $200 \text{ days}$ . If a source initially contains pure  $^{210}_{83}\text{Bi}$ , after how long will the rate of  $\alpha$ -particle emission reach a maximum? Derive any equations you use showing your working. You may assume a trial solution of the form  $Ae^{-\lambda_1 t} + Be^{-\lambda_2 t}$ . [10]
- e) If a sample of  $1 \text{ gram}$  of  $^{210}\text{Po}$  exists at the present time, estimate the number of atoms of Po at the present time and after  $100 \text{ days}$ . What is the mass of the Polonium sample after  $100 \text{ days}$ ? [5]
- f) Given the relatively short half life of  $^{210}\text{Po}$  state (without calculation) whether or not you would expect the  $Q$  of its alpha decay to be large or small giving your reason. [2]

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B2)

- a) Explain, in terms of nuclear binding energy, why fusion processes are generally energetically more favourable for light nuclei whereas fission reactions are only important for heavy nuclei. [2]
- b)  $^{235}_{92}\text{U}$  is an  $\alpha$  emitter with a half life of  $10^9$  y. It is also able to undergo spontaneous fission with a half life of  $10^{16}$  y. What is meant by the term spontaneous fission? [2]
- c) Calculate the Coulomb potential energy for  $^{235}_{92}\text{U}$  briefly existing in the form of:
- An alpha particle and an isotope of  $_{90}\text{Th}$ .
  - A  $^{117}_{46}\text{Pd}$  nucleus and a  $^{118}_{46}\text{Pd}$  nucleus.

By sketching the shape of the nuclear potential as a function of radial distance  $r$ , explain which mechanism is responsible for spontaneous fission and why this is suppressed for large mass fragments. [6]

- d) Fission may be induced by the collision of a neutron with the Uranium nucleus.  $^{235}_{92}\text{U}$  has a large cross section for absorption of thermal neutrons leaving the new uranium isotope in an excited state with a mass of  $219883.5 \text{ MeV}/c^2$ . Calculate the excitation energy. Explain what is meant by the activation energy for the fission of  $^{236}_{92}\text{U}$ . If the activation energy is found to be  $6.2 \text{ MeV}$  what is the minimum kinetic energy of the incident neutron for fission to occur? [3]
- e) What are the typical kinetic energies of thermal and fast neutrons? Draw a diagram showing how the following cross sections vary with incident neutron energy
- $^{235}_{92}\text{U}(n,f)$
  - $^{235}_{92}\text{U}(n,\gamma)$
  - $^{238}_{92}\text{U}(n,f)$

Use this to explain what the process of moderation is and why it is a critical aspect in the design of a nuclear reactor. [10]

- f) The deuteron has a mass of  $2.014102 \text{ u}$ . Calculate its binding energy in MeV. Explain why the fusion reaction  $^1\text{H} + ^1\text{H} \rightarrow ^2\text{He}$  does not occur and so direct proton fusion can only proceed through the reaction  $^1\text{H} + ^1\text{H} \rightarrow ^2\text{H} + e^+ + \nu_e$ . In the solar fusion cycle why is this step known as a bottleneck? The hydrogen burning fusion process can occur through three distinct routes (in the absence of carbon), each involving one of the following partial steps:

- $^1\text{H} + ^1\text{H} \rightarrow ^2\text{H} + e^+ + \nu_e$
- $^7\text{Be} + e^- \rightarrow ^7\text{Li} + \nu_e$
- $^8\text{B} \rightarrow ^8\text{Be} + e^+ + \nu_e$

Which earth based observations allow researchers to distinguish reaction (ii) from the others? [7]

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B3)

- a) Explain what is meant by the Extreme Independent Particle Model. [2]
- b) Motivate the inclusion of the pairing term in the semi-empirical mass formula, describing its behaviour for even-even, odd-odd and even-odd nuclei. Stable nuclei tend to have  $Z=N$  for low  $A$  and  $N>Z$  for large  $A$ . Explain the physical origin of this tendency. [5]
- c) Sketch a graph of the behaviour of the neutron absorption cross section as a function of the neutron number. Comment on any notable features. [4]
- d) Determine the proton and neutron shell configurations for the following nuclei as well as their expected spin and parity assignments for the ground states.
- i.  ${}^7_3\text{Li}$
  - ii.  ${}^{23}_{11}\text{Na}$
  - iii.  ${}^{33}_{16}\text{S}$  [6]

- e) The nuclear magnetic moment  $\mu$  for an unpaired proton can be calculated in the Independent Particle Model as

$$\text{i. } \langle \mu \rangle = \left[ g_l \frac{j(j+\frac{3}{2})}{j+1} - \frac{1}{2} \frac{j}{j+1} g_s \right] \mu_N \quad \text{for } j=l-1/2$$

$$\text{ii. } \langle \mu \rangle = \left[ g_l \left( j - \frac{1}{2} \right) + \frac{1}{2} g_s \right] \mu_N \quad \text{for } j=l+1/2$$

Where  $g_l = 1$  and  $g_s = +5.58$  (in units of  $\mu_N$ ) for a free proton. What is the prediction of this model for the magnetic moment of  ${}^{209}_{83}\text{Bi}$ ? Why does this prediction differ from the measured value of  $4.1\mu_N$ ? [6]

- f) The average of the neutron absorption cross section for heavy nuclei is approximately 0.1 barn. Would you expect this cross section for  ${}^{209}_{83}\text{Bi}$  to be significantly different from the value of 0.1 barn? [2]
- g) The nucleus  ${}^{42}_{21}\text{Sc}$  has an odd number of neutrons and protons. What are the range of ground state spin values that this nucleus could have? [5]

End of Examination Paper

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